A MAP estimation algorithm with Gibbs prior using an IIR filter

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Abstract. The MAP method is a wide spread technique used in many

signal processing problems, e.g., image restoration, denoising and 3D reconstruction. When there is a large number of variables to estimate, the MAP method leads to an huge set of linear or non-linear equations which must be numerically solved using time consuming algorithms. This paper proposes a fast method to compute the MAP estimates in large scale problems, based on the solution of a linear or linearized set of equations by low pass filtering the ML solution. A family of space varying IIR filters is derived from the MAP criterion with data dependent coefficients. Other filter expressions can be derived by the same approach, using different assumptions about the prior or other filter design strategies. The filter approach proposed in this paper is much faster than the calssic solution and provides additional insights about the structure of

Experimental results are provided to assess the performance of the proposed methods with Gaussian and non Gaussian noise models.

1 Introduction

the problem.

The maximum likelihood (ML) method is widely used to estimate signals from noisy data. In the image processing field, for instance, the ML method is often used to solve problems of image restoration [1], denoising [2], deblurring [3] and 2D or 3D reconstruction [4].

The number of publications on medical imaging using statistical approaches [5], e.g., the ML method, received much more attention after the work of Shepp and Vardi [7] on emission tomography.

However, the estimated solutions, e.g. in reconstruction problems, tend to be noisy and the estimation algorithms tend to converge slowly in the vicinity of the solution [8]. This is often due to the ill-conditioned nature of the estimation problem. In the case of 2D or 3D reconstruction the lack of data is the main difficulty [10], since there is often less then one observation per unknown.

To overcome these difficulties the MAP method is often used since it regularizes the ML solution by smoothing it. By using a prior distribution containing a prior knowledge about the unknowns variables, the MAP method introduces

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additional constraints on the solution, improving ill-conditioned nature of the problem. It has been shown, by several authors, that this approach overcomes the difficulties caused by the lack of data, producing a smoothing effect by interpolating the data. Furthermore, the smoothing also reduces the amount of noise present in the ML estimates [9,6].

Gibbs distribution have been traditionally used in MAP problems. This class of priors leads to simple formulations of the estimation problem. The equivalence with the Markov random fields allows to easily obtain the joint probability distribution from a set of local distributions [11].

The Gibbs prior with quadratic potential functions are simple leading to a simplified formulation of the MAP method. Furthermore, the recursive structure of the resulting equations can be used to speed up the estimation algorithms.

In this paper the problem of MAP estimation of signals if formulated as a filtering process. It is shown that the MAP estimate using a Gibbs prior with quadratic potential functions, can be computed by filtering the ML estimation with two IIR filters: a causal IIR filter and an anti-causal IIR filter, leading to computational gains up to one order of magnitude. These filters are first order IIR filters with space varying impulse response. It is shown, that the cutoff frequencies of the proposed filters are adaptatively adjusted according to the number of observations and some sufficient statistics.

In this paper, two models are considered: additive Gaussian noise and multiplicative modeled by a Rayleigh distribution. These are typical models which can be used in a wide range of applications.

The formulation presented here can be applied to problems, like image restoration, involving one observation per unknown or to problems, like 3D reconstruction, in which some unknown variables are not observed.

The approach described in this paper can be extended to derive other Gibbs priors corresponding to higher order filters, more selective, performing better at transitions. In fact, with this filtering approach it is clear why the MAP estimation method using Gibbs prior with quadratic potential functions, does not perform well in the transitions. Under this perspective, the algorithm is filtering the ML with a first order filter, which in general tend to blur the abrupt transitions.

2 Problem Formulation

Let $X = \{x_i\}$ be a sequence of N unknown variable to be estimated and $Y = \{y_i\}$ a sequence of observations. Each element of Y, y_n is, itself, a set of n_i observations of x_i . In typical problems of image restoration $n_i = 1$, which means that there is one observation per pixel. On the contrary, in 3D reconstruction, the number of observations per voxel varies from voxel to voxel. For instance, in free-hand 3D ultrasound the number of observations associated to non inspected voxels is zero $(n_i = 0)$. On the contrary if given voxel is intersected by several cross sections $n_i > 1$.

In this paper the MAP method is used to estimate X from the observations Y. This method estimates X by maximizing an energy function,

$$\hat{X} = \arg\min_{U} E(Y, X) \tag{1}$$

where

$$E(X,Y) = -l(X,Y) - \log p(X) \tag{2}$$

 $l(X,Y) = \log(p(Y|X))$ is the likelihood function, and p(X) is the a prior distribution associated to the unknown variables.

For sake of simplicity let us assume that y_i is normal distributed (later we will consider other distributions), with $p(y_i) = N(x_i, \sigma^2)$ corresponding to the following observation model

$$y_i = x_i + w_i \tag{3}$$

with $p(w_i) = N(0, \sigma^2)$. If the observations are independent, the log-likelihood function is given by

$$l = C - \frac{\beta}{2} \sum_{i,k} (y_{ik} - x_i)^2 \tag{4}$$

where $\beta = 1/\sigma^2$ and y_{ik} is the kth observation of the unknown x_i .

The prior p(X) used in (2) plays an important role in the estimation process when there is lack of information about the variables X (n_i small), since the ML estimates are very poor in this case [10].

In this paper we will consider that p(x) is a Gibbs distribution with quadratic potential functions [13, 11]. This is equivalent to assuming that the vector X is a Markov random field [12, 11]. Therefore

$$p(X) = \frac{1}{Z} e^{\sum_{i} V_i(X)} \tag{5}$$

where Z is the partition function and $V_i(X)$ is the potential function associated to the i-th unknown. Assuming that X is a 1D signal and assuming that each variable x_i has two neighbors, x_{i-1}, x_{i+1} ,

$$p(X) = \frac{1}{Z} e^{-\frac{\alpha}{2} \sum_{i} (x_i - x_{i-1})^2}$$
 (6)

The parameter α defines the strength of the links among neighbors and it is pre-defined. Therefore, the energy function to be minimized is

$$E(Y,X) = \frac{\beta}{2} \sum_{i,k} (y_{i_k} - x_i)^2 + \frac{\alpha}{2} \sum_i (x_i - x_{i-1})^2$$
 (7)

The constants C and Z were discarded because they do not contribute to the solution.

3 Optimization

To minimize (7) the following stationary conditions must be met

$$\frac{\partial E(Y,X)}{\partial x_i} = 0 \tag{8}$$

which lead to the following set of linear equations

$$x_i = (1 - k_i)x_i^{ML} + k_i\bar{x}_i \qquad i = 1, ..., N$$
 (9)

where x_i^{ML} is the maximum likelihood estimation of x_i , k_i is a parameter that depends on the data and \bar{x}_i is the average value of the neighbors of x_i

$$x_i^{ML} = \frac{1}{n_i} \sum_{k=1}^{n_i} y_{i_k}^2 \tag{10}$$

$$k_i = \frac{1}{1 + \frac{\beta n_i}{2\alpha}} \tag{11}$$

$$\bar{x}_i = \frac{x_{i-1} + x_{i+1}}{2} \tag{12}$$

To minimize the border effects it is assumed that $x_0 = x_2$ and $x_{N+1} = x_{N-1}$. Taking this into account, (9) can be written as follows

$$Ax = b \tag{13}$$

where

and

$$b = [(1 - k_1)x_1^{ML}, (1 - k_2)x_2^{ML}, ..., (1 - k_N)x_N^{ML}]^T$$
(14)

The estimation of (13) amounts to the solution of a linear system of equations which can be performed by using either iterative (e.g. Gauss elimination method) or non iterative methods (e.g. Gauss-Seidel method). Since the number of unknowns is often very large (e.g. on the order of a million) iterative methods are preferred since they provide an approximate solution with acceptable computational effort. In order to obtain an efficient solution the structure of A can be considered.

In the next section we will show that the system (13) can be solved using two space varying IIR filters, obtained from (9).

4 IIR filter

Equation (9)

$$x_i = (1 - k_i)x_i^{ML} + \frac{k_i}{2}(x_{n-1} + x_{n+1})$$
(15)

defines a non causal recursive filter [17] with x_i^{ML} as input. Assuming that k_i is constant the filter impulsive response, g_i , can be computed. The general form is

$$g_i = Ca^{|i|} \tag{16}$$

where C and a are computed replacing (16) in (15) and making $x_i^{ML} = \delta(i)$

$$\begin{cases} Ca^0 = (1-k) + \frac{k}{2}(Ca^{|-1|} + Ca^{|1|}) & i = 0 \\ Ca^i = \frac{k}{2}(Ca^{i-1} + Ca^{i+1}) & i \neq 0 \end{cases}$$

$$\begin{cases} C = \frac{1-k}{1-ak} \\ a^2 - \frac{2}{k}a + 1 = 0 \end{cases}$$

leading to

$$\begin{cases} a = \frac{1 \pm \sqrt{1 - k^2}}{k} \\ C = \frac{1 - k}{-(\pm)\sqrt{1 - k^2}} \end{cases}$$

Since $0 \le k \le 1$ and C > 0, only one solution is feasible, i.e.,

$$\begin{cases} a = \frac{1 - \sqrt{1 - k^2}}{k} \\ C = \frac{1 - k}{\sqrt{1 - k^2}} \end{cases}$$

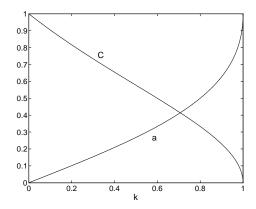


Fig. 1. Impulse response parameters.

Fig.1 shows the dependence of a and C on $k \in [0,1]$. As it can be observed, a is monotonic increasing with k and limited to the interval [0,1].

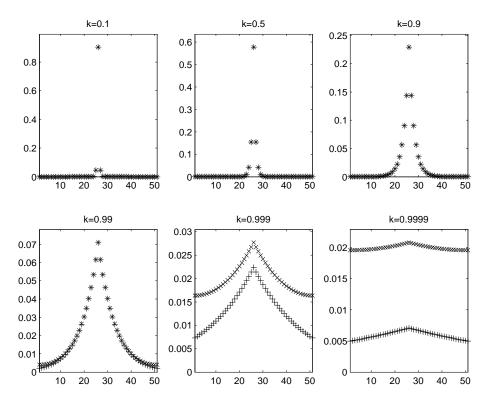


Fig. 2. Impulsive response computed from (13) (x) and from (16) (+).

Fig.2 shows the impulsive response (16) and the solution of (13) for several values of k. As it can be seen, both solutions are identical, except for k close to 1 (k > 0.99). These differences are due to border effects. In fact, equation (15) is not valid at n = 1 and n = N.

The impulsive response defined by (16) can not be used in a recursive way because it is not wedge supported [17]. Therefore, we will decompose it as a sum of two wedge supported impulsive responses, one causal and other anti-causal, i.e., one depending only on past inputs and outputs and other depending only on future inputs and outputs [18]. Therefore,

$$g_n = g_n^+ + g_n^- (17)$$

where

$$g_n^+ = \begin{cases} Ca^n & n > 0 \\ \frac{C}{2} & n = 0 \\ 0 & n < 0 \end{cases}$$
$$g_n^- = \begin{cases} 0 & n > 0 \\ \frac{C}{2} & n = 0 \\ Ca^{-n} & n < 0 \end{cases}$$

$$g_n^- = \begin{cases} 0 & n > 0\\ \frac{C}{2} & n = 0\\ Ca^{-n} & n < 0 \end{cases}$$

where it was assumed $g_0^+ = g_0^-$ to impose symmetry. Applying the Z transform to the previous equation we obtain

$$G(Z) = G(Z)^{+} + G(Z)^{-}$$
(18)

$$G(Z)^{+} = \frac{C}{2} \frac{1 + aZ^{-1}}{1 - aZ^{-1}}$$
 (19)

$$G(Z)^{-} = \frac{C}{2} \frac{1 + aZ}{1 - aZ} \tag{20}$$

The solution of (13) is the sum of two terms

$$x_i = x_i^+ + x_i^- (21)$$

where

$$x_i^+ = g_i^+ * x_i^{ML} = \frac{C_i}{2} (x_i^{ML} + a_i x_{i-1}^{ML}) + a_i x_{i-1}$$
 (22)

is a causal space varying recursive filter and

$$\bar{x}_{i} = g_{i}^{-} * \bar{x}_{i}^{ML} = \frac{C_{i}}{2} (\bar{x}_{i}^{ML} + a_{i} \bar{x}_{i+1}^{ML}) + a_{i} \bar{x}_{i+1}$$
 (23)

is an anti-causal space varying recursive filter where

$$\begin{cases} a_i = \frac{1 - \sqrt{1 - k_i^2}}{k_i} \\ C = \frac{1 - k_i}{\sqrt{1 - k_i^2}} \end{cases}$$

The MAP estimates defined in (1) can be obtained as follows. First the maximum likelihood estimates X^{ML} is computed. Then the ML estimate are filtered with a causal filter $G(Z)^+$ and with an anti-causal filter $G(Z)^-$. The solution is obtain by adding both results.

5 Frequency analysis

It is now clear that the regularization imposed by the prior is equivalent to filtering the ML estimates with a first order low-pass filter that smoothes the transitions, reducing the noise present in the maximum likelihood estimation.

The low-pass filters (19) and (20) present a 0.5 gain at d.c., a pole located at a and a zero at -a (see Fig.3). The cutoff frequency, depending on the pole position, depends on the data and on the parameter α of the prior as it can be observed on (4) and (11)¹(see Figs.3,4). Therefore, the following conclusions can be stated:

1) since $0 \le a \le 1$, the pole of the first order filter is always inside the unit circle and the filter is always stable.

¹ Note that a(k) is monotonic with k.

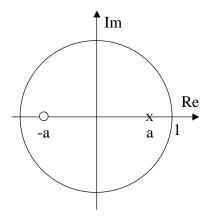


Fig. 3. Pole and zero position of G(z).

- 2) the parameter k (parameter a) (see eq.(11)) decreases with the number of data points n_i , i.e., the bandwidth of the filter increases with the amount of available data and decreases as the number of data points goes to zero. The algorithm compensates the lack of confidence in the data by decreasing the filter bandwidth (see Fig.4).
- 3) the bandwidth of the filter decreases when the regularization parameter α increases and when the variance of the data, $\sigma^2(x) = 1/\beta$, increases.
- 4) the algorithm is implemented in such way that when there is no data, $n_i = 0$, k = 1 (a = 1). In this case $x_i = \bar{x}_i$, i.e., the estimate depends only on the average value of the neighbors.

Until this point we have been working with an additive Gaussian noise. However, the method can be used with other noise models. For instance, the Rayleigh distribution is often used to model the multiplicative noise present in signals obtained using coherent radiation, e.g., LASER [14], SAAR [16] or ultrasound [15]. In a previous work, the authors have derived expressions similar to (15) in the context of 3D ultrasound by using a Rayleigh model for the multiplicative speckle noise present in the ultrasound images [21]. In the case of the Rayleigh observations (15) is still valid with

$$\begin{cases} x_i^{ML} = \frac{1}{2n_i} \sum_k (y_i^k)^2 \\ k_i = \frac{1}{1 + \frac{n_i}{2\alpha(x_i^{ML})^2}} \end{cases}$$

In this case k_i depends on the number of observations, n_i , as before, but it also depends on the ML estimate. Therefore an additional property is valid.

5) the bandwidth decreases with the increase of x_i^{ML} , i.e., as stronger regularization is applied for large values of x_i^{ML} than for smaller values (the noise amplitude is larger in high intensity regions). This behaviour is a consequence of the multiplicative type of noise associated to the Rayleigh distribution.

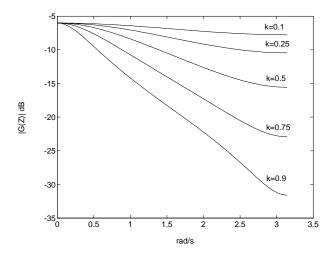


Fig. 4. Bode diagrams of G(Z) for several different values of k.

As shown, the MAP estimation problem can be interpreted as a space varying filtering process. Adopting a Gibbs prior with a quadratic potential the MAP estimation process can be implemented by using two first order IIR filters. This approach can also be used to derive other type of filters associated, for instance, to higher order Gibbs priors, which allow the improvement of the estimation performance at transitions (see [19]).

In the next section two examples of application are presented using synthetic and real data.

6 Experimental Results

Experimental tests were performed to evaluate the algorithm in 1D and 2D signal restoration, with synthetic and real data.

Each problem is solved using the standard MAP method and the fast algorithm based on space-varying IIR filters proposed in this paper. Examples with Gaussian and non-Gaussian (Rayleigh) noise are considered.

6.1 Synthetic data

Let us consider a synthetic signal defined as follows,

$$x_i = \begin{cases} 150 & 50 \le i \le 100 \\ 50 & otherwise \end{cases}$$

The observation vector Y is obtained by adding Gaussian white noise $\eta_i = N(0, 20^2)$ to each sample (see Fig.5).

The MAP solution was computed by both methods, i.e, by solving (13) and by low pass filtering using (21).

Both solutions are displayed in Fig.5. Since both curves coincide they can not be distinguish. To minimize the border effects, the unknowns x_0 and x_{N+1} , used in equations (22) and (23) respectively were defined as follows

$$x_0 = \frac{1}{2w} \sum_{i=1}^{w} x_i^{ML} \tag{24}$$

$$x_{N+1} = \frac{1}{2w} \sum_{i=N-w}^{N} x_i^{ML} \tag{25}$$

i.e., x_0 and x_{N+1} are initialized with half of the average value of the ML estimates inside window with length w = 5.

The two solutions are not identical. However, their difference is so small that can not be observed in Fig.5. The two solutions are almost identical within the interval [0, 150], except in the vicinity of the boundaries.

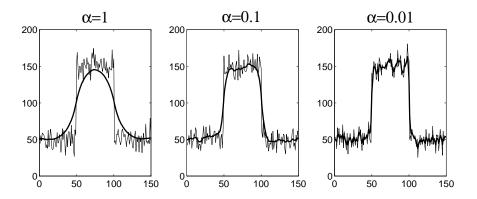


Fig. 5. Synthetic data: Noisy and filtered data for $\alpha = 1, 0.1, 0.001$.

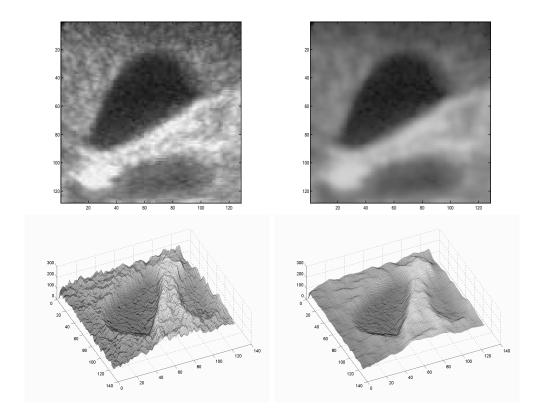
6.2 Ultrasound image

This example considers the problem of noise reduction in ultrasound images, using a multiplicative model for the noise (Rayleigh model). The MAP estimates of the original image was computed by both methods, i.e., by solving the linear set of equations (13), obtained by linearization of the non linear cost function (2), and by using the IIR filters defined in (21). X^{ML} and k are computed using (5).

We have used the following separable filter to process the ultrasound image

$$G(Z_1, Z_2)_{2D} = G(Z_1)G(Z_2)$$
(26)

where G(Z) is given by (18). Separable filters allow fast filtering procedures based on two steps: in the first step the filter G(Z) is applied to each column of the ultrasound image and, in a second step G(Z) is applied to each row of the image obtained in the previous step. Fig.6 shows an ultrasound image (left), and the MAP estimates obtained by the IIR filter (right). The results achieved by solving (13) are nor shown since they are similar. Fig.7 shows the matrix of coefficients



 ${f Fig.\,6.}$ First column:ultrasound image, Second column: MAP estimates .

 k_i . As noted before, the Rayleigh distribution, corresponding to a multiplicative type of noise, make the coefficients depend not only on the amount of data, but also on the data itself (see (5)). Therefore, the lighter zones of Fig.7 correspond to regions where the cutoff frequency of the IIR filter is smaller and consequently the regularization effect imposed by the prior is higher. On the contrary, in the darker regions, corresponding to the darker regions on the original ultrasound image, the regularization effect imposed by the prior is smaller.

Fig.8 shows the image obtained by computing the absolute difference of the images computed by the matrix inversion method and by the filtering method. The difference is small, except at the transitions and at the borders. The signal

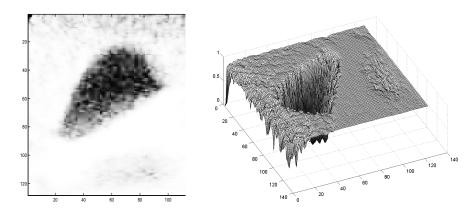


Fig. 7. Matrix of coefficients k_i .

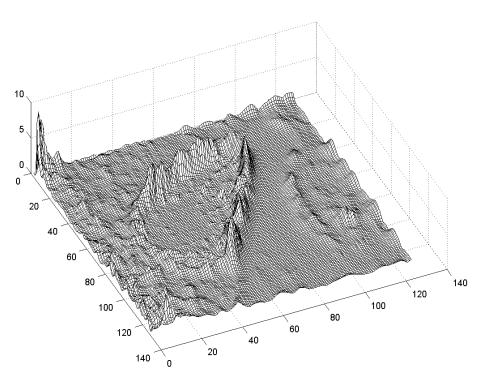


Fig. 8. Error image (SNR = 43.64dB).

do noise ratio is $SNR = S_A - S_\Delta = 43.64dB$. S_A is the energy of the image X_A estimated using equation (13) and computed as $S_A = 10 \log_{10}(X_A.X'_A)$. $S_\Delta = 10 \log_{10}[(X_A - X_F)(X_A - X_F)']$ is the energy of the error image displayed in Fig.8. X_F is the image estimated using equation (21). The largest difference between X_A and X_F is observed at the origin due borders effect (as expected).

7 Conclusions

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